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Title:

From Basic Research to Homeland Security: Muon Radiography and Large-Area Detectors

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Institute for Nuclear and Particle Astrophysics and Cosmology meeting, San Diego, October 5



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ABSTRACT

From Basic Research to Homeland Security: Muon Radiography and Large-Area Detectors

William C. Priedhorsky and Richard C. Schirato

Talk at meeting of Institute for Nuclear and Particle Astrophysics and Cosmology San Diego, CA 3 October 2003

There are a number of connections between detector technology as developed for nuclear and particle astrophysics, and detection needs for homeland security. We set out two of them. We begin by discussing the basic detection requirements, and the emissions from uranium and plutonium. One way to detect the neutron flux is to develop a very large neutron detector, using technology developed for neutrino physics. Another way to detect dense, high-Z material is by its scattering of natural, cosmic-ray neutrinos, which are present all the time. We present concepts and simulations for the former, and experimental results and simulations for the latter.

From Basic Research to Homeland Security: Muon Radiography and Large-Area Detectors

October 3-5, 2003 INPAC meeting, San Diego

William Priedhorsky and Richard Schirato Los Alamos National Laboratory

on behalf of the muon radiography team:

K. N. Borozdin, G. E. Hogan, C. L. Morris, A. Saunders, L. J. Schultz, M. E. Teasdale

and the VLAND/neutrino team:

Martin Cooper, Gerry Garvey, Bill Louis, Shawn McKenney, Geoff Mills, Dick Mischke, Joel Moss, Richard Schirato, Richard Van de Water, Nate Walbridge, Hywel White; UNM: Doug Fields + students

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Los Alamos

Homeland Security Needs

- Detection of Special Nuclear Materials (SNM) and nuclear WMD's
- Detection of Radiation Dispersal Devices (RDD's) or illicit radioisotopes
- Cargo/vehicle inspection for other threats, e.g. chemical explosives, drugs, etc.



Signatures: Spontaneous Fission

Nuclide	Specific Intensity [n/(g-s)]
235	0.0003
238 U	0.0136
²³⁸ Pu	2590.
²³⁹ Pu	0.022
²⁴⁰ Pu	1020.
²⁴¹ Pu	0.05
²⁴² Pu	1720.
²³⁷ Np	0.0001

Implication: U doesn't really emit neutrons, and the few percent of ²⁴⁰Pu in reactor-produced material is dominant

Solution: Induce fissions with neutrons or gammas



Signatures: Gamma emission

239
Pu $\xrightarrow{24,000 \text{ y}}$ 235 U* + α

Relatively complex decay scheme; major gamma rays:

129.28 keV Intensity 140,000 γ /(g-s)

413.69 keV Intensity 34,000 γ /(g-s)

$$^{235}U$$
 $^{710,000,000 \text{ y}}$ $^{231}\text{Th}^* + \alpha$

Much simpler decay scheme; major gamma ray:

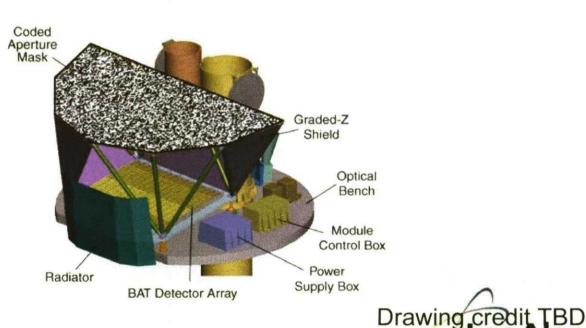
185.72 keV Intensity 43,000 γ /(g-s)



Technology flow: "Homeland Security" to Basic Science

- DOD and DOE have historically funded detector research which has found applications in Basic Science
 - e.g. organic and inorganic scintillators, light detectors, Monte Carlo simulations, etc.
- Recent Example:
 - CdZnTe room temperature gamma/X-ray detectors

Burst Alert Telescope on SWIFT Mission



Technology flow: Basic Science to "Homeland Security"

Possibilities from INPAC-related research include:

- Large area, efficient neutron and gamma detectors
- Cosmic ray muon radiography/interrogation
- Other detector technologies discussed at this meeting...



LANSCE Radiation Detection Portal

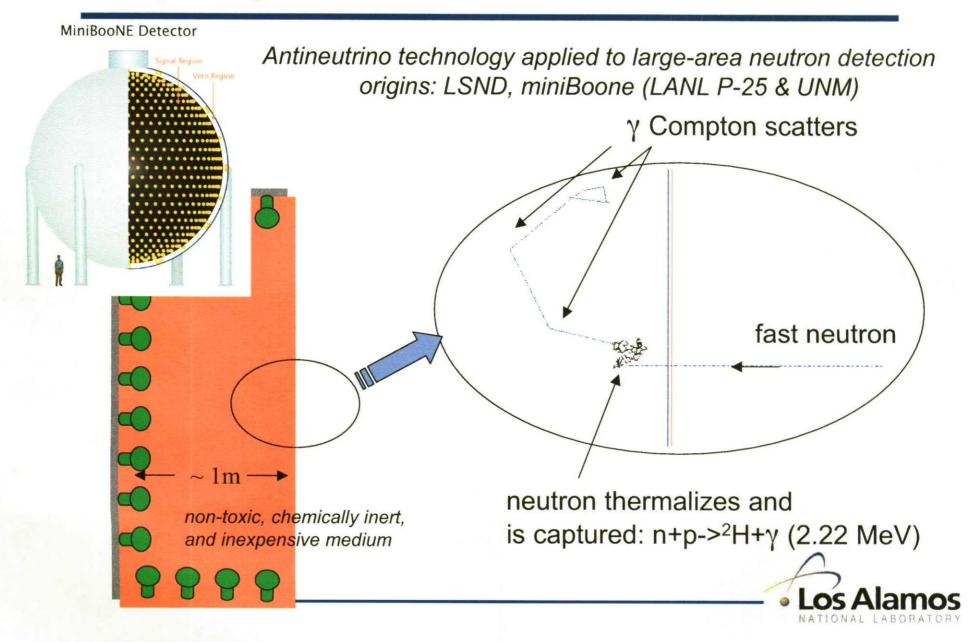
- 40 year-old technology
- Facility-specific
 - solid plastic scintillators for γ detection (Compton)
 - ³He proportional counters for neutron detection
 - hand-held monitors (NaI) for cargo searches

2 x 6"x31"x1.5" plastic scintillators
2 x 2"x40" moderated ³He prop. counters





Very Large Area Neutron Detector- VLAND



VLAND as a full-energy gamma ray detector

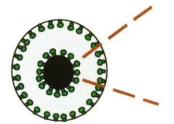
Full energy containment: modest isotopic sensitivity [15%/√E (MeV)] unlike current plastic

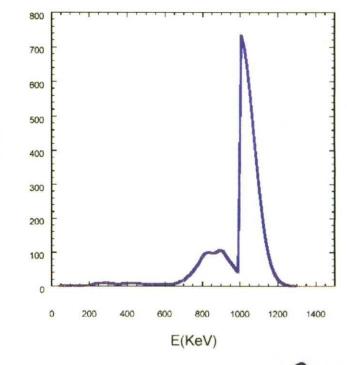
Counts

scintillator portal monitors

▶ better resolution possible KAMLAND: 7.5%/√E

directional sensitivity





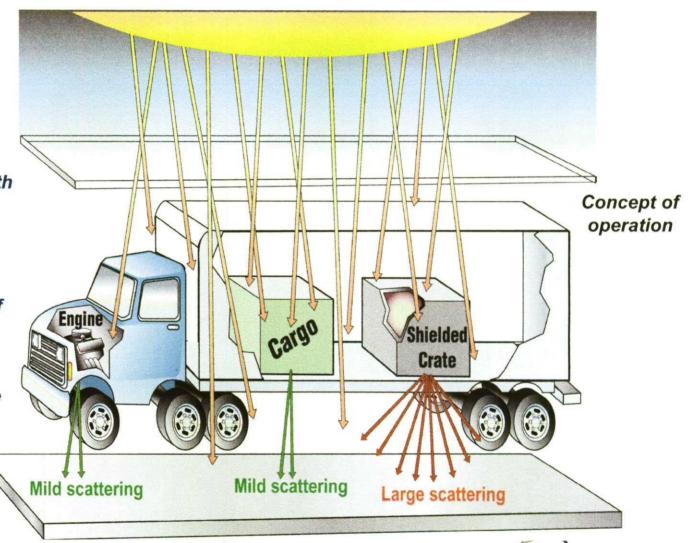


Zero-dose radiography

Objective:
penetrating radiography with
no artificial dose

Application:
Prevent illicit movement of nuclear materials

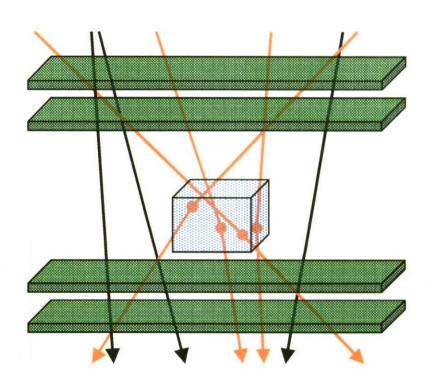
The heavier the shield, the easier the detection





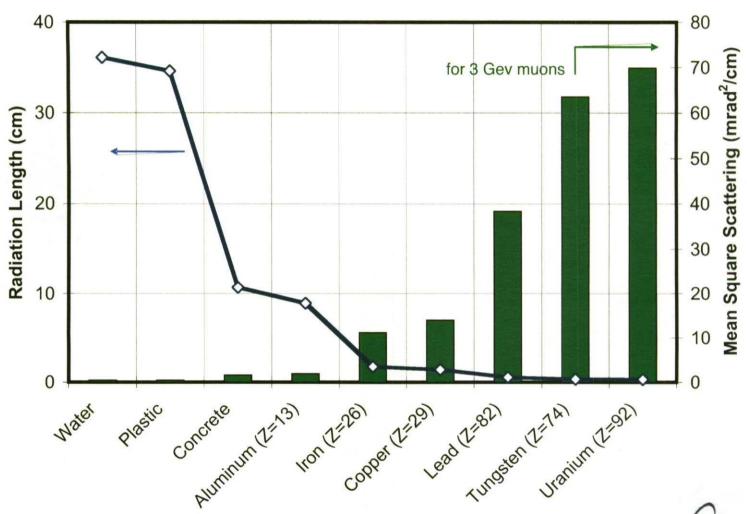
Basic Concept of Muon Radiography

- Track <u>individual</u> muons (possible due to modest event rate).
- Track muons <u>into</u> and <u>out of</u> an object volume.
- Determine scattering angle of each muon.
- Infer material density within volume from data provided by many muons.
- New tomographic reconstruction algorithms are required.



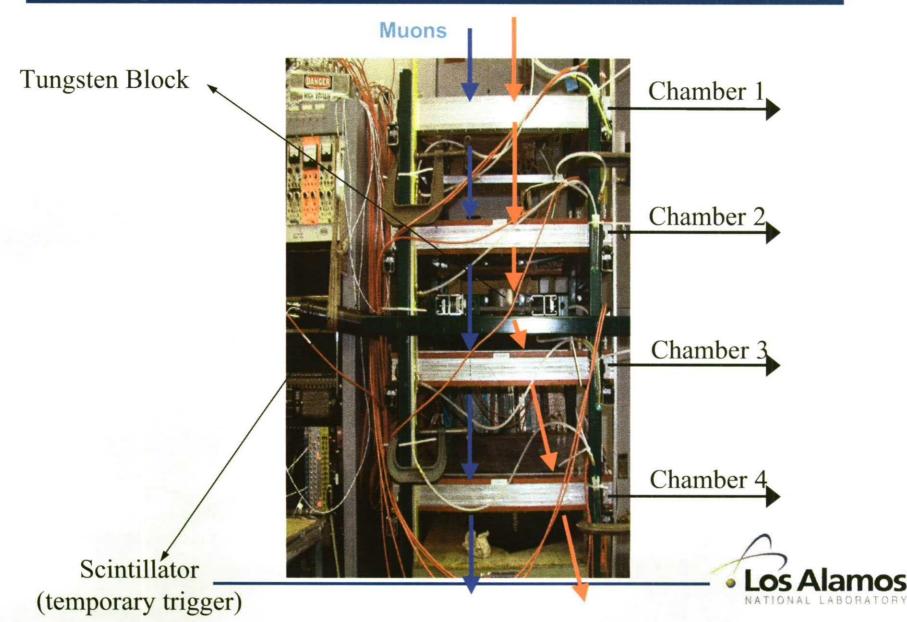


Scattering is Material Dependent

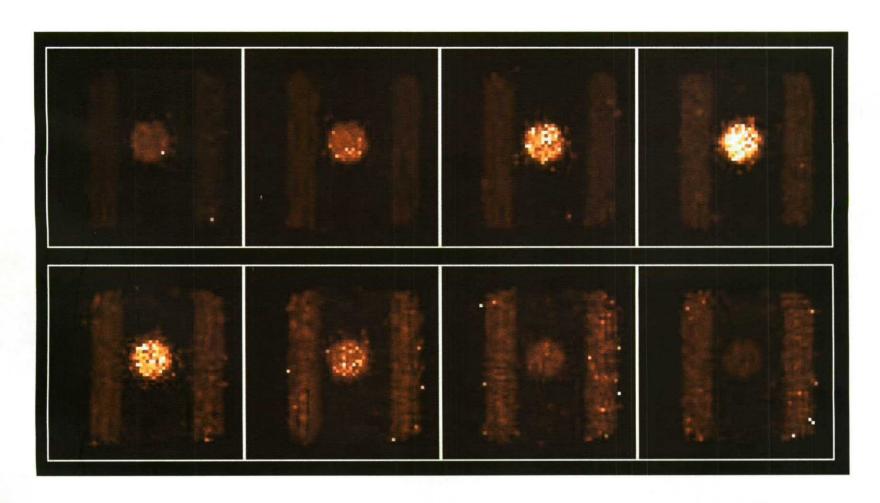




Prototype instrument



Slices through reconstructed volume





Experiment/Simulation Comparison

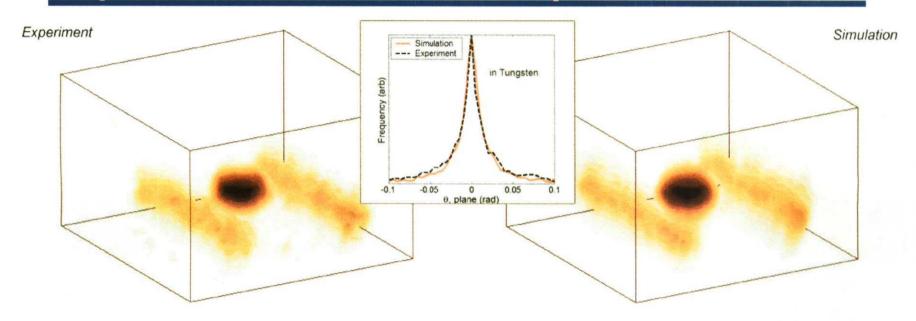


Image of a 5 cm high, 10 cm diameter tungsten block, and steel supporting struts, equivalent exposure 30 minutes

Simulation validated in this experiment, shows promise of nuclear material detection

Nature, 20 March 2003

Visibility in the press:

http://www.nature.com/nsu/030317/030317-7.html

http://news.bbc.co.uk/2/hi/science/nature/2868041.stm

http://www.guardian.co.uk/uslatest/story/0,1282,-2492519,00.html

http://www.space.com/businesstechnology/technology/nuclear_detection_030319.html

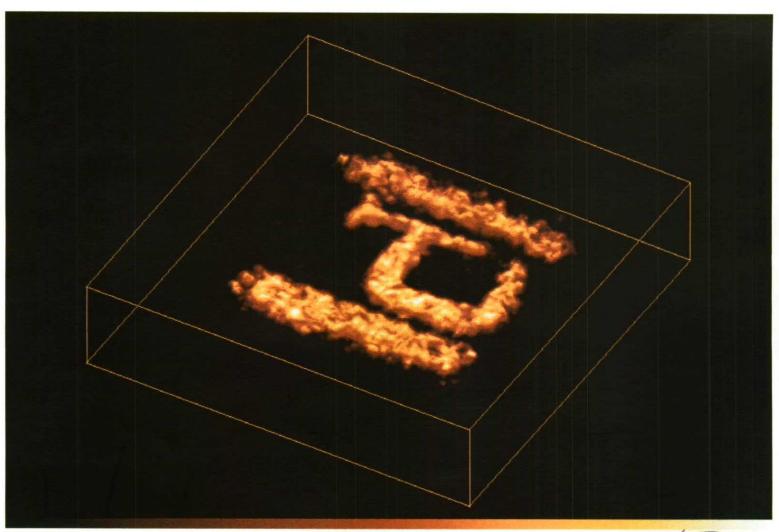
http://www.physicsweb.org/article/news/7/3/11

http://news.nationalgeographic.com/news/2003/03/0319 030319 cosmicrays.html

http://www.newscientist.com/news/news.jsp?id=ns99993526

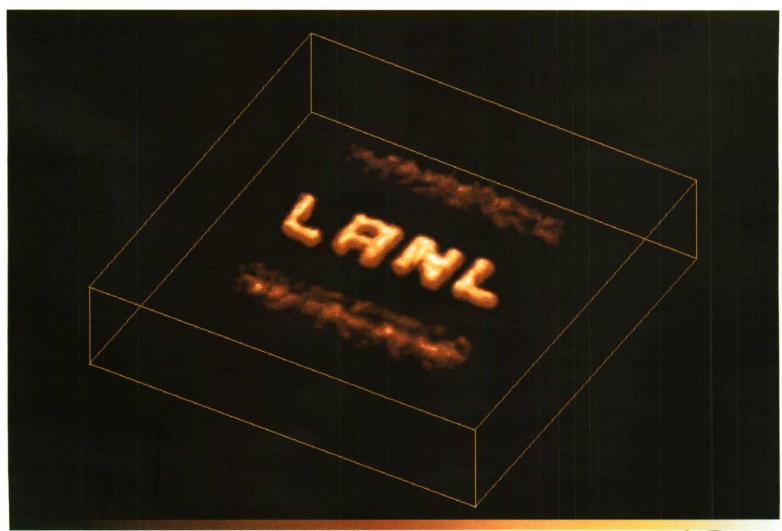


Clamp in 3D



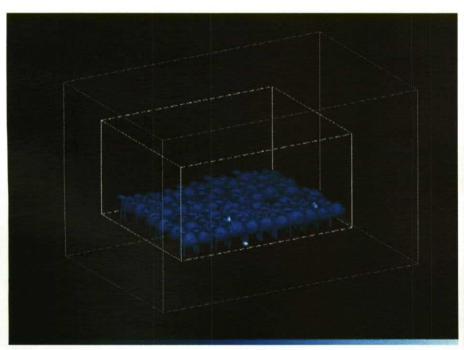


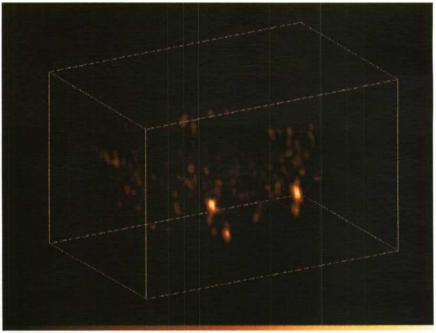
Perspective view





Uranium objects in a shipping container





OBJECTS

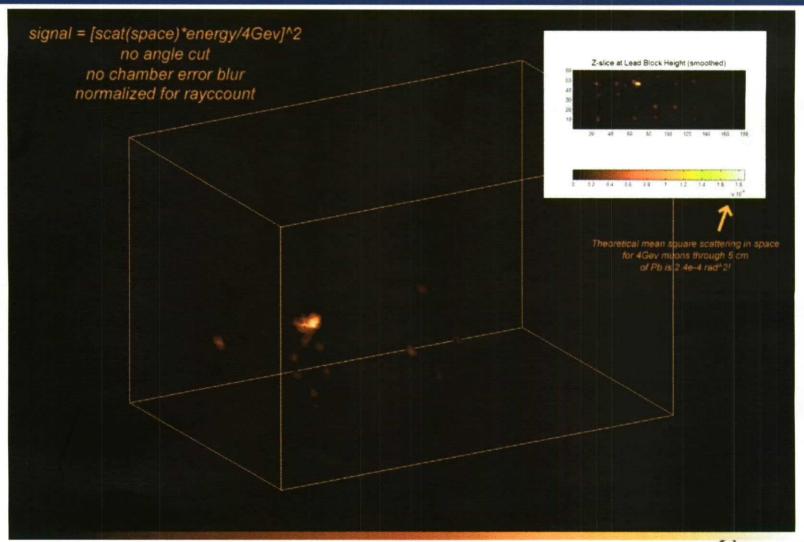
2.4 m x 2.4 m x 6 m container 3 mm steel wall 9 cm x 9 cm x 12 cm uranium "pigs" (60 cm)³ water "sheep"

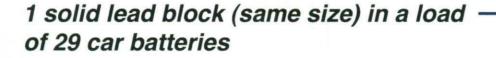
SIMULATED RADIOGRAPH

1 minute exposure
90% detection probability
>>6 sigma rejection of false alarms
"point of closest approach" reconstruction



Car batteries vs. solid lead block





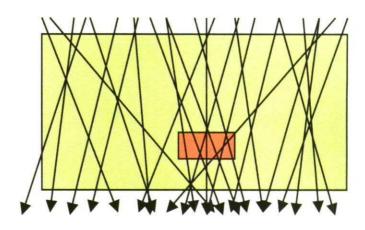


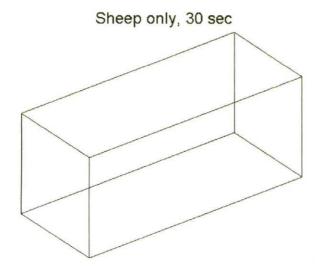
Clutter Rejection: Ray Crossing Algorithm

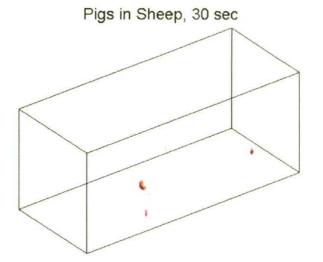
Apply a bend angle cut, keep only high bending rays

Apply an approach cut, keep only rays which approach one another closely.

Create 3-D histogram of approach locations.







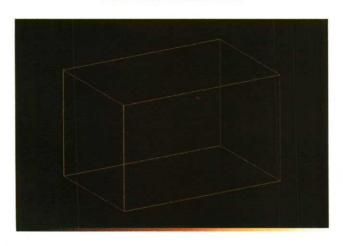


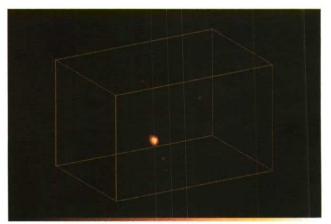
A Scene with more Clutter

No contraband

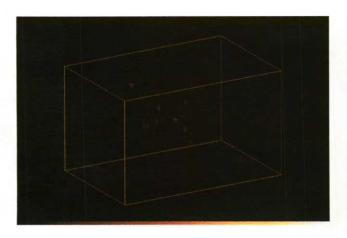
With 3 pigs

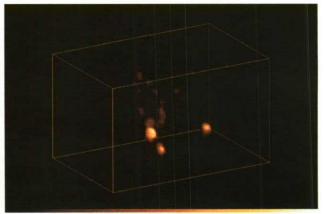
30 second exposure





120 second exposure





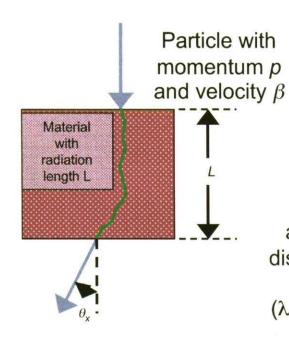
10 tons of distributed iron filling the container



Backup slides



Physics of interaction: multiple scattering



Scattering distribution is approximately Gaussian

$$\frac{dN}{d\theta_x} = \frac{1}{\sqrt{2\pi}\theta_0} e^{-\frac{\theta_x^2}{2\theta_0^2}}$$

and the width of the distribution is related to $\theta_0 = \frac{13.5}{p\beta} \sqrt{\frac{L}{\lambda}}$ (λ is a radiation length)

$$\theta_0 = \frac{13.5}{p\beta} \sqrt{\frac{L}{\lambda}}$$

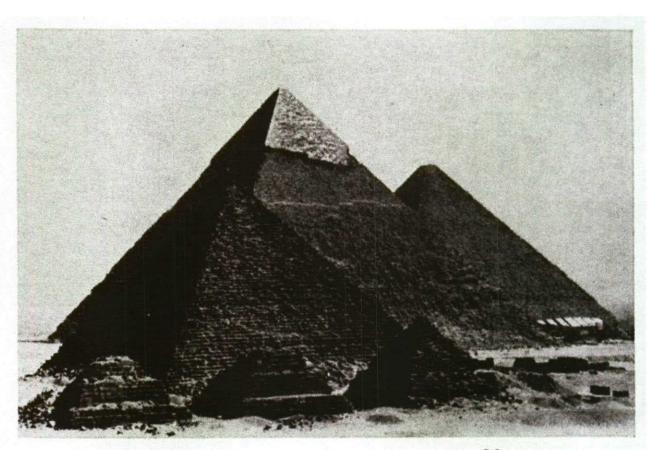
Scattered particles carry information from which material may be identified.

Material	λ,cm	θ_0 , mrad*
Water	36	2.3
Iron	1.76	11.1
Lead	.56	20.1
*10 cm of	f material, 3 G	ev muons



A previous target

Fig. 1 (top right). The pyramids at Giza. From left to right, the Third Pyramid of Mycerinus, the Second Pyramid of Chephren, the Great Pyramid of Cheops. [© National Geographic Society]



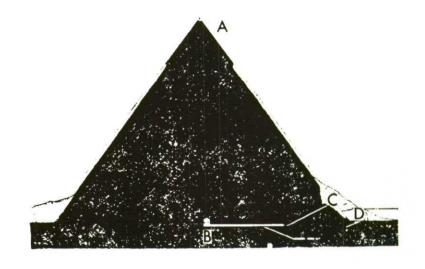


Muon mapping of Chephren's Pyramid Science, 167, 832 (1970)

Search for Hidden Chambers in the Pyramids

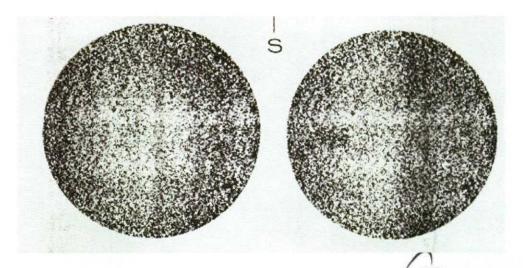
The structure of the Second Pyramid of Giza is determined by cosmic-ray absorption.

Luis W. Alvarez, Jared A. Anderson, F. El Bedwei, James Burkhard, Ahmed Fakhry, Adib Girgis, Amr Goneid, Fikhry Hassan, Dennis Iverson, Gerald Lynch, Zenab Miligy, Ali Hilmy Moussa, Mohammed-Sharkawi, Lauren Yazolino

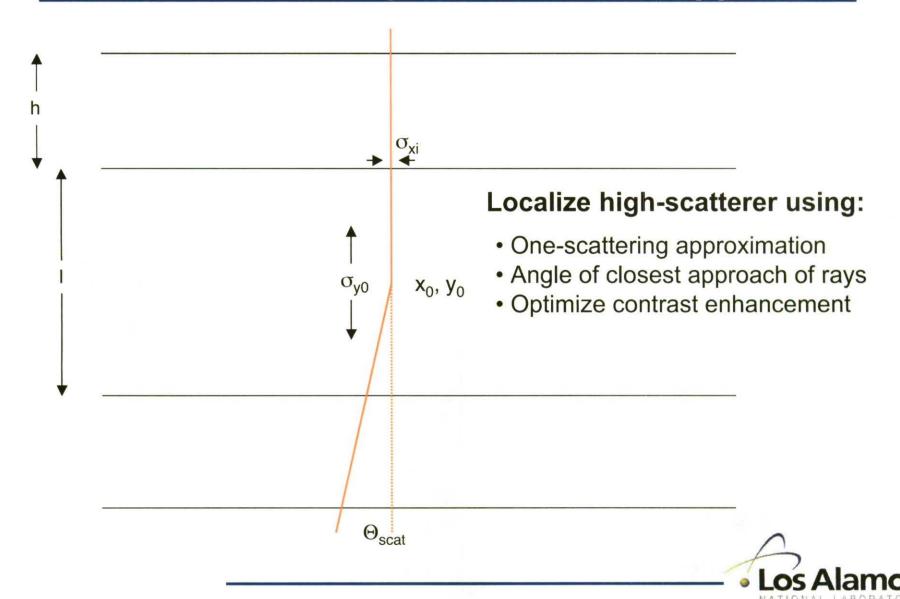


left: actual image
right: simulated image,
with hidden
chamber
as in Cheops

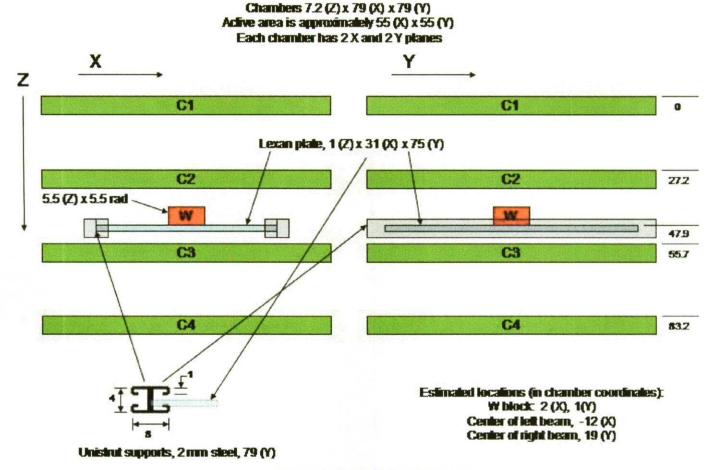
Only absorption exploited in this analysis – not scattering



Reconstruction via point of closest approach



Test object: tungsten cylinder



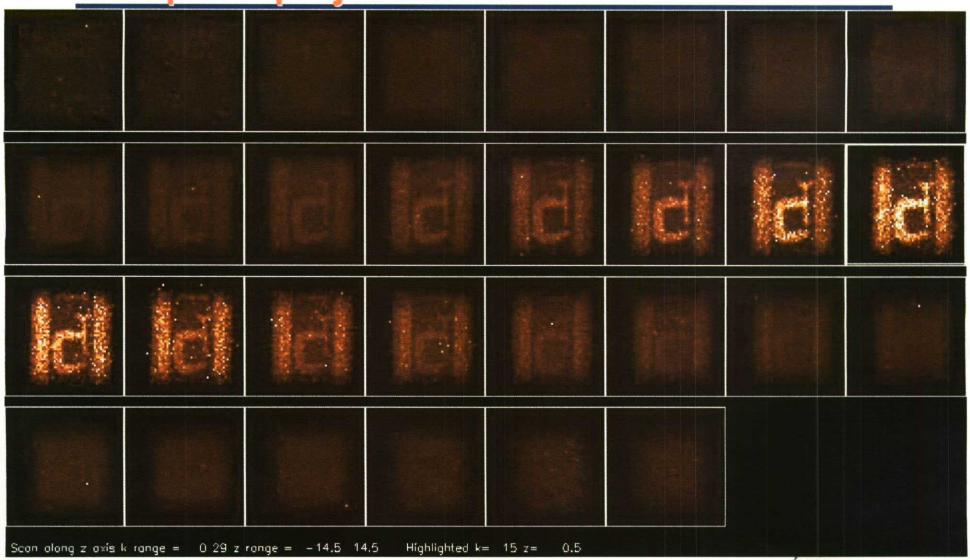




Another object

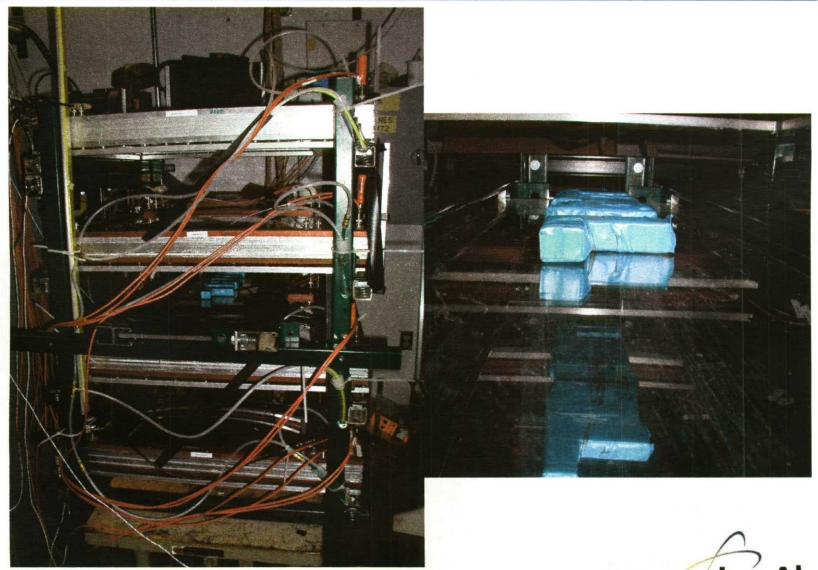


Clamp in z-projections





One more object





... and its reconstructed slice

